

# Quantum gravity: a phenomenological perspective

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# Quantum gravity: a phenomenological perspective

- Naive consequences of QG :
  - Minimal length or "size",
  - QFT and its domain of validity,
  - Maximal acceleration & force,
  - Maximal energy density & pressure,
  - Holographic principle,
  - Energy-dependent  $\hbar$ .
- A tentative geometrization: covariant mass bound conjectures
- Remarks on Minimal Length Scale scenarios, GUP, and vacuum energy

# Fundamental constants and their possible meaning

- Intuitively, quantum gravitational world is built on fundamental length, time, and energy scales (of order of the Planck units, say)
- In this point of view, the emergence of  $c$ ,  $h$ ,  $G$  is non trivial and linked to fundamental kinematics/dynamics of the QG world.

$$c = \frac{l_p}{t_p}, \quad G = \frac{l_p^3}{m_p t_p^2}, \quad \hbar = \frac{m_p l_p^2}{t_p},$$

- Most intuitive example : causality in QG give raise to an universal limiting speed  $c = l_p/t_p$ .
- What are the other physical effects leading to the emergence of  $G$  and  $h$  ?

# Fundamental constants and their possible meaning

- Standard papers :  $\hbar$  is structural, as is  $c$ , because leading to a whole new description of physics ; but  $G$  is somehow « only » a coupling constant
- I actually disagree with this : Schwarzschild result  $G m / c^2 < 1/2$  is one of the main unexpected result of GR, and its main non trivial consequence, pretty much like the  $E = mc^2$
- $G/c^2$  is structural constant as well, in the sense that  $m < 1/c^2 / 2G$
- Non intuitive aspect : not a direct bound on the density, but on the lineic mass which is bounded from above
- On the other hand  $\hbar$  is less intuitive

# Basic inequalities

- **Naive QG** : consider a system of mass  $m$  (or energy  $E$ ), of "size"  $l$ , observed during  $t$ .
- Then we have
  - $l/t < c$
  - $G m / l c^2 < 1/2$
  - $\sigma = m l^2/t > \hbar$  (quantum of action)
- **Problems** :
  - What do we mean by size of a physical system (outside spherical symmetry)?
  - Schwarzschild only apply in sph. symmetry
  - Not very clear what  $t$  refers to.

# Basic inequalities

- Assume for now

**Law 1.** *Existence of a maximal speed. Locally, and in any frame, the speed  $v$  of physical systems must satisfy:*

$$v \leq c. \quad (2)$$

**Law 2.** *(First formulation). Existence of a mass bound. In  $D = 4$  spacetimes, any physical system with typical size  $l$  has a maximal mass  $m$  given by:*

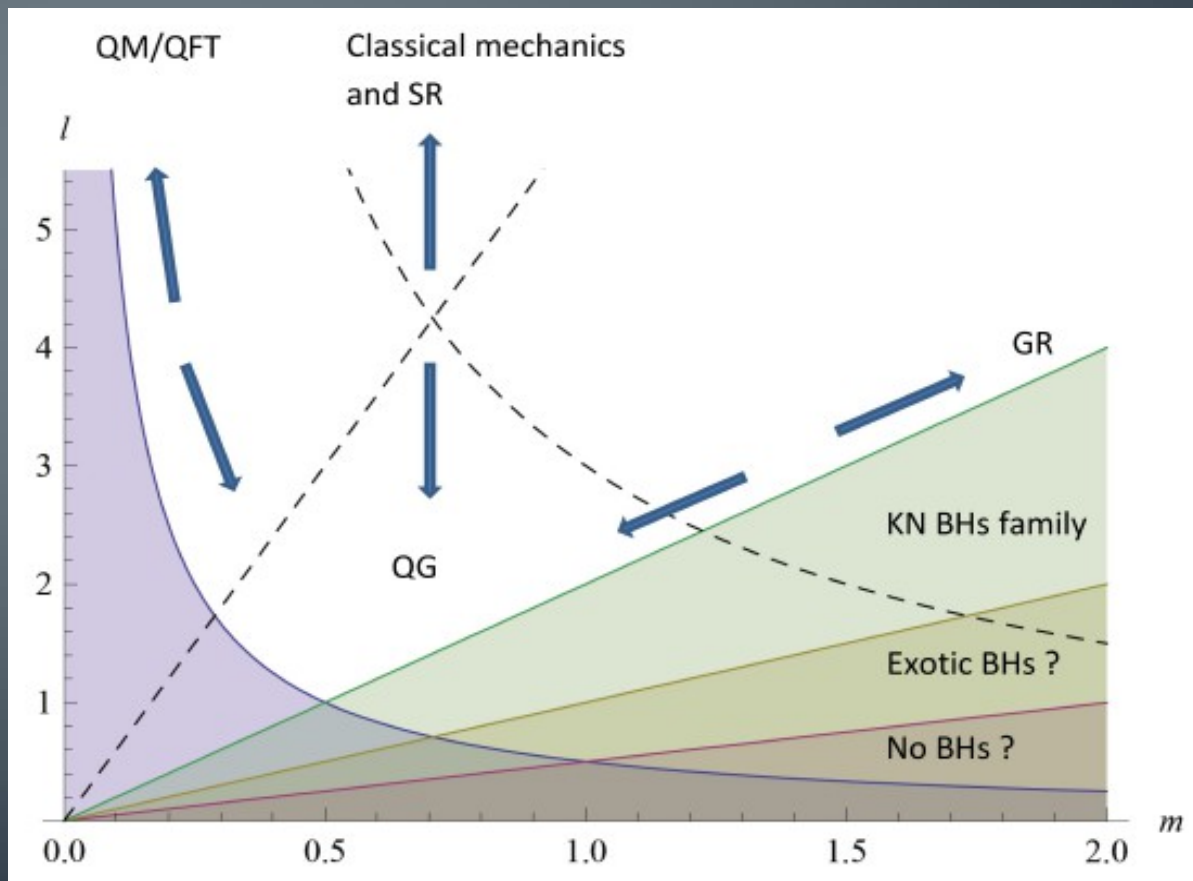
$$\frac{Gm}{lc^2} \leq \mathcal{O}(1). \quad (4)$$

**Law 3.** *Existence of the quantum of action. There exists a positive pure number  $\beta$  such that any physical system of size  $l$  and mass  $m$  satisfies*

$$ml \geq \frac{\beta \hbar}{c}. \quad (8)$$

# Consequence 1 : minimal "size"

- $ml > 1$  and  $m < 1$  and thus  $l > \text{Max}(m, 1/m) > 1$  : minimal "size" of physical systems or minimal uncertainty  $\Delta x > 1$
- But no bounds on the mass  $m$ .



## ... But no minimal length!

- Consider now a box  $L$  filled with one quanta  $p=\hbar/\lambda$
- Use  $m/L < 1/2$  and  $m=E/c^2=\hbar/\lambda$
- Get  $1/L < \lambda < L$  : subplanckian wavelength/transplanckian frequencies allowed (ie without the box collapsing into a BH)
- In strong disagreement with the basic idea of DSR (Doubly Special Relativity with an invariant Planck length)
- No paradox : the box must be larger than Planck length, but can be filled with a subplanckian quanta.
- Physically because the quanta is delocalized, and Einstein eqs. care only about the energy density. Density can be subplanckian while quanta's momenta is transplanckian.



# Csq 2: Holographic properties

- Consider a system of mass  $m$  made of  $N$  dof, with  $m = N \omega$
- Use prev. result :  $1/l < \lambda < l \Rightarrow 1/l < \omega < l$

- Thus get

$$\frac{m}{l} < N < ml$$

- Right hand side similar to Bekenstein bound

- $$N \sim S \leq 2\pi ER$$

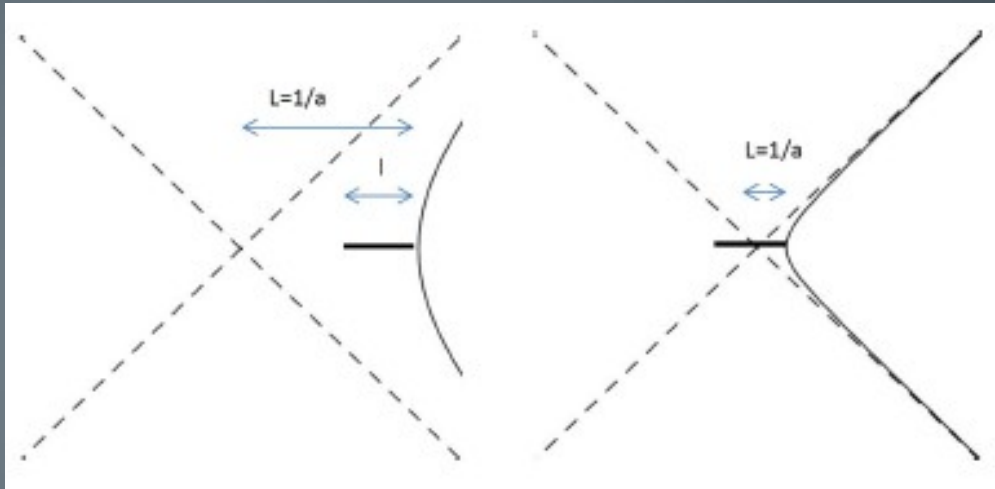
- And using  $m < l$ , we get the holographic principle in its naive form

$$\mathcal{O}(1) < N < l^2$$

- Remark, entropy bound saturated requires  $m \sim l$  (hence heavy  $m > 1$  systems), but  $N \ll l^2$  when  $m \ll l$ ; and  $N \sim 1$  iff  $m \sim 1/l$ : systems on Compton line have no substructures.
- (Later generalized to  $S < A/4$ , covariant entropy bound, Bousso)

# Csq 3: maximal acceleration

- Objects have finite size + special relativity :  $a < 1/l$



Moreover  $l > 1$ , thus  $a < 1$  bounded by (the huge) Planckian acceleration

# Csq 3: maximal acceleration and force

- But there may exist a much tighter bound !

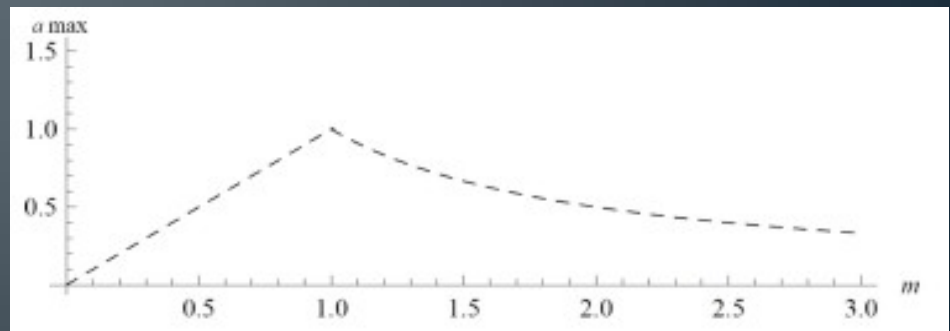
$$\begin{cases} a < \frac{1}{l} < m < 1 & \text{if } m < 1 \\ a < \frac{1}{l} < \frac{1}{m} < 1 & \text{if } m > 1 \end{cases} \quad (17)$$

- Insert back fundamental constants, and get

$$\begin{cases} a < \frac{mc^3}{\hbar} & \text{if } m < 1 \\ a < \frac{c^4}{Gm} & \text{if } m > 1 \end{cases}$$

-> Caianiello's max acc (1980)

-> New ?



# Csq 4: maximal force and power

- Similarly, mass-dependent maximal force

$$\begin{cases} f < \frac{m^2 c^3}{\hbar} & \text{if } m < 1 \\ f < \frac{c^4}{G} & \text{if } m > 1 \end{cases} \quad (23)$$

- And power, with  $P_{\text{max}} = f_{\text{max}} c$
- For both acceleration, force, and power, the weird thing is that maximal allowed values in the deep  $m \ll 1$  regime or  $m \gg 1$  regime are only quantum relativistic ( $\hbar, c$ ) or respectively general relativistic ( $G, c$ ) !!
- Needs more thinking and modelling !

# Tentative geometrization

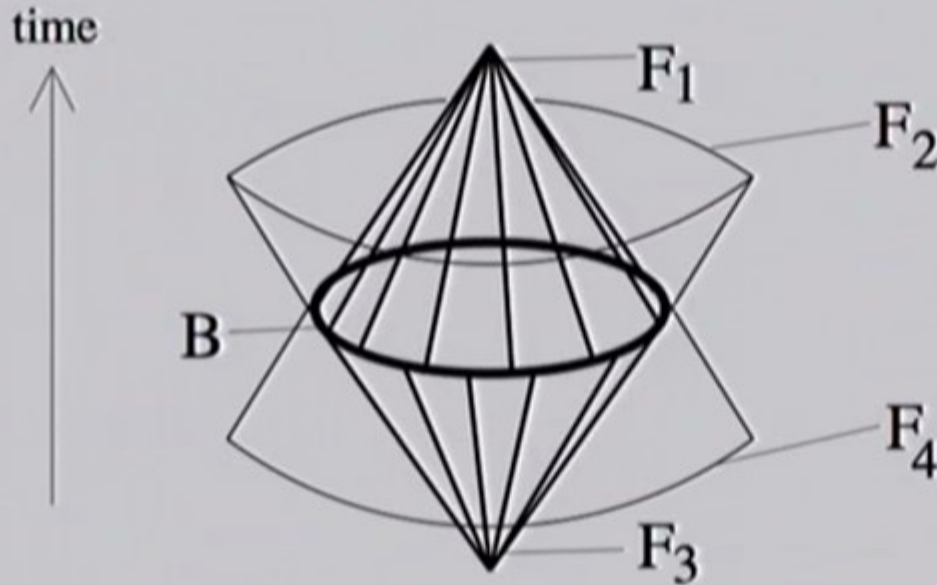
Back to our naive inequalities. What is the size of system ? Example, of a spaghetti ? We need to go beyond Schwarzschild result, ie beyond a spherical symmetric mass bound. Also, how do we get a covariant bound ?

Entropy bound debate (Bekenstein and many others)

-> R. Bousso in 98/99 « covariant entropy bound » :

# Tentative geometrization

## Covariant Entropy Bound



For any two-dimensional surface  $B$  of area  $A$ , one can construct lightlike hypersurfaces called light-sheets. **The total matter entropy on any light-sheet is less than  $A/4$  units:  $S \leq A/4G\hbar$ .**



# Tentative geometrization

So why not a covariant mass bound ?

**Law 2.** *(Second formulation). Covariant mass bound conjecture. In a four-dimensional spacetime which is free of horizons, and given any surface  $B$  of codimension 2, the integrated mass on the light-sheets  $L[B]$  must satisfy:*

$$m[L[B]]^2 < \frac{A[B]}{16\pi G^2}, \quad (5)$$

(Tricky part : requires a local definition of mass)

Plausibility ? If a spacetime has some static event horizons, then its mass satisfies :

$$16\pi m^2 \geq A_T$$

This is the **Penrose inequality/conjecture** ; and the covariant mass bound is somehow nothing more than its converse

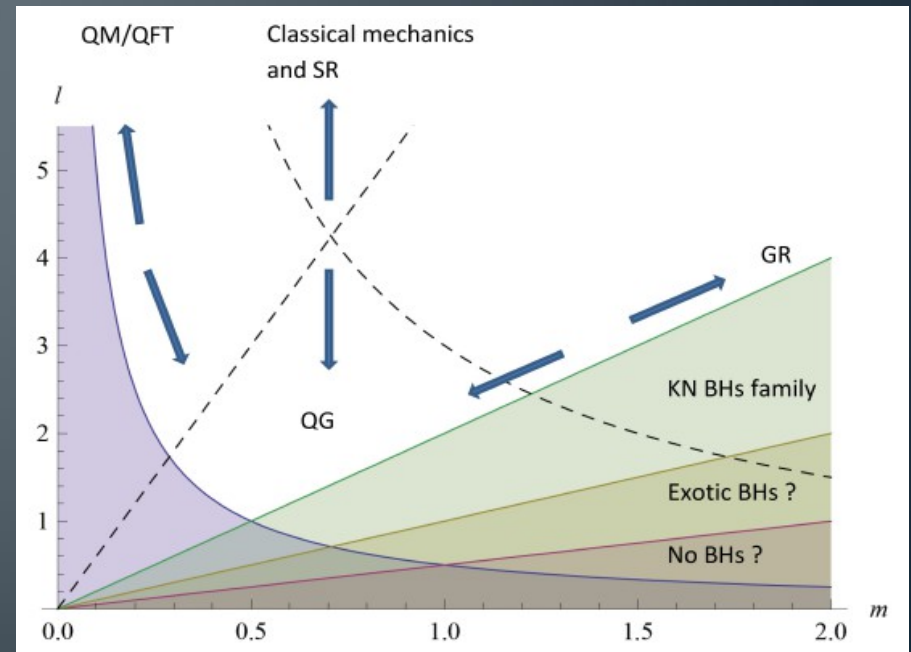
# Tentative geometrization

Example : Kerr Newmann Black Hole family has

$$A_H = 4\pi \left( 2m^2 - q^2 + 2m \left( m^2 - q^2 - J^2/m^2 \right)^{1/2} \right),$$

And thus :  $4\pi m^2 \leq A_H \leq 16\pi m^2$

In terms of  $l = \text{Sqrt}(A/4/\text{Pi})$ ,  
(areal radius)





# And the quantum of action?

In fact squaring the Compton inequality  $\lambda > \hbar/mc$ , and  $\lambda \rightarrow \lambda^2 \rightarrow A$   
One may write, in 4D, the quantum of action as :

$$A > \frac{4\pi\beta^2\hbar^2}{m^2c^2},$$

(For some pure number beta)

To be compared to the gravitational inequality

$$A > \frac{16\pi G^2 m^2}{c^4}.$$

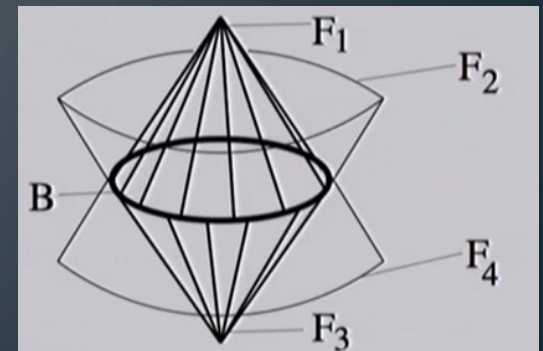
Where one notes the remarkable symmetry

$m \rightarrow 1/m$

Note : we don't have a minimal size  $\lambda$ ,

But rather a minimal area in fact

**Interpretation ?**



# Csq: Geometrized version of QG inequalities

Example

Mass/area relations :  $4 \pi \hbar^2 / A < m^2 < A / (16 \pi G^2)$

Acceleration :  $a < 1/l$  may become  $a^2 < 1/A$  (Rovelli 2013 in LQG)

Maximal density and pressure  $\rho = m/l^3 < 1/l^2$  becomes  $\rho < G/A$

etc

# Phenomenological models

DSR ? -> maybe not relevant

Minimal uncertainty in position space ? -> Generalized Uncertainty Principle (GUP)

$$\Delta x \Delta p > f(\Delta p^2),$$

where the standard GUP is  $f(x) = 1 + \beta x$

Modification of canonical commutation relation -> modif of QM axioms

Max acceleration ? Modification of inertial physics like

$$-d\tau^2 = dx_\mu dx^\mu + \frac{\hbar^2}{m^2 c^2} d\dot{x}_\mu d\dot{x}^\mu, ,$$

Non trivial question : effective models of QG, or a new starting point ?

# Vacuum energy

Why it shall affect vacuum energy ?

Very roughly (QM) :  $H = p^2/2m + m \omega^2 X^2/2$

+ Heisenberg principle :  $\Delta x \Delta p > 1/2 \Rightarrow H \sim p^2 + 1/p^2$

Compute its minimum and find  $H_{\min} = \hbar \omega/2 \Rightarrow$  vacuum energy

But with the same Hamiltonian and the GUP, it obviously changes

It gives  $p^2 = m\omega/\text{Sqrt}[4 + b^2 m^2 \omega^2] \rightarrow \text{cste}$  at large  $\omega$  ;

And thus  $H_{\min} \sim f(\omega) \sim \text{cst} + O(\omega^2)$  at large  $\omega$  : even more UV divergent !

But requires a careful identification of  $H$ , and a rigorous GUP-QFT calculation

# « Conclusions »

Phenomenological approaches to QG in the literature often based on thought experiments with explicit or implicit use of spherical symmetry (eg. Measurement's precision limited by the formation of a horizon)

However we need to go beyond this symmetry to find more robust mass bounds, and thus « model independent » consequences of QG.

⇒ Nice covariant mass bounds for both grav. and quantum properties. The area naturally appears as the relevant quantity !

⇒ Surprising bounds which are not all  $(h,c,G)$  dependent, but only  $(h,c)$  or  $(G,c)$  (though all coming from the fact that minimal size exists)

Model building : what is the point ? Is it only an effective description, or a whole new start ? eg. Max acc and modif of minkowski space